



Exciton Liquid in Coupled Quantum Wells

Michael Stern *et al.*

Science **343**, 55 (2014);

DOI: 10.1126/science.1243409

This copy is for your personal, non-commercial use only.

If you wish to distribute this article to others, you can order high-quality copies for your colleagues, clients, or customers by [clicking here](#).

Permission to republish or repurpose articles or portions of articles can be obtained by following the guidelines [here](#).

The following resources related to this article are available online at www.sciencemag.org (this information is current as of January 7, 2014):

Updated information and services, including high-resolution figures, can be found in the online version of this article at:

<http://www.sciencemag.org/content/343/6166/55.full.html>

Supporting Online Material can be found at:

<http://www.sciencemag.org/content/suppl/2014/01/02/343.6166.55.DC1.html>

Exciton Liquid in Coupled Quantum Wells

Michael Stern,*† Vladimir Umansky, Israel Bar-Joseph

Excitons in semiconductors may form correlated phases at low temperatures. We report the observation of an exciton liquid in gallium arsenide/aluminum gallium arsenide-coupled quantum wells. Above a critical density and below a critical temperature, the photogenerated electrons and holes separate into two phases: an electron-hole plasma and an exciton liquid, with a clear sharp boundary between them. The two phases are characterized by distinct photoluminescence spectra and by different electrical conductance. The liquid phase is formed by the repulsive interaction between the dipolar excitons and exhibits a short-range order, which is manifested in the photoluminescence line shape.

More than 40 years ago, it was shown theoretically that excitons in semiconductors may form a condensate at cryogenic temperatures (1). Subsequently, a wealth of classical and quantum phases, ranging from electron-hole plasma to superfluid liquid, have been predicted to occur in this material system (2–4). Spatially indirect dipolar excitons with the electrons and holes residing in separate coupled quantum wells (CQWs) offer a promising test bed for these phenomena (5–8). They are characterized by a strong repulsive dipole-dipole interaction, which can be easily engineered by designing the sample structure. Their density, n , and the distance between the centers of the wells, d , are key parameters that govern the formation of condensed states in this system. Together they set the strength of the dipole interaction with the surrounding carriers. The interaction energy, $\delta E(n) = \iint V(r)n(r)d^2r$, can be expressed in terms of an effective radius, r_0 , around every exciton, within which the carrier density is depleted as a consequence of the repulsive interaction V

$$\delta E(n) = \frac{\langle n \rangle e^2}{\epsilon} \left[\sqrt{r_0^2 + d^2} - r_0 \right] \quad (1)$$

Here, $\langle n \rangle$ is the average density, e is the electron charge, and ϵ is the dielectric constant. $\delta E(n)$ can be directly measured in a photoluminescence experiment as an energy shift of the recombination energy with density.

Theory predicts that at densities well above the Mott transition, the system may form an exciton liquid (3, 4). For d smaller than the exciton Bohr radius a_B , the exchange interaction can compensate for the dipole repulsion, and a quantum liquid is formed. For $d \geq a_B$ this phase is not stable, but the system may still form a classical liquid. The origin of this classical liquid is related to the formation of a region around each exciton,

with radius $\rho \approx 4d^2/a_B$, where the wavefunctions of other excitons are exponentially small (4). When $n\rho^2 \geq 1$, the wave functions of neighboring excitons do not overlap, and quantum correlations are suppressed. Classical correlations, on the other hand, are very strong and may give rise to short-range order. This classical limit of strongly interacting dipoles is the focus of our work.

The system we studied was formed by two gallium arsenide (GaAs) quantum wells with well widths of 12 and 18 nm, separated by a 3-nm $\text{Al}_{0.28}\text{GaAs}$ barrier (energy diagram shown in Fig. 1A); the distance between the centers of the wells is thus $d = 18$ nm, almost twice as large as a_B . The carriers were photogenerated by a laser and separated by the application of a gate voltage V_g in the direction perpendicular to the wells. More experimental details are given in (9).

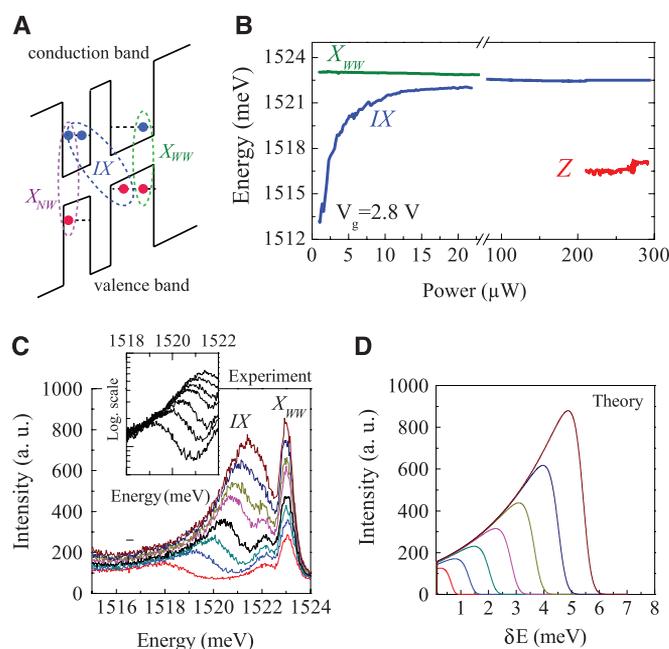
In Fig. 1B, we show the evolution of the spectrum with increasing laser power P . Between

1 and 100 μW , the indirect recombination energy (IX) increases monotonically with P as the density n increases. This shift to higher energy is the manifestation of the growing interaction energy, $\delta E(n)$, and it allows us to determine the electron-hole density: For example, at 20 μW , $\delta E = 9\text{meV}$, which corresponds to $\sim 2 \times 10^{10}$ to $3 \times 10^{10} \text{cm}^{-2}$. δE is not linear with power and saturates as the IX peak approaches the wide-well exciton line, X_{WW} . This behavior can be understood when it is noted that the recombination rate Γ is proportional to the overlap integral of the electron and hole wavefunctions (IO): As the density increases, the potential drop between the quantum wells is screened, and Γ increases exponentially. An interesting implication of this behavior is that at $P \gg 20 \mu\text{W}$, the density across the sample is almost constant and depends only weakly on the local power.

In Fig. 1C, we show the spectra at various powers, where the blueshift of the IX peak with power is clearly evident. The IX spectra at all powers are characterized by a pronounced exponential tail at the low-energy side (II). This line shape reflects the broad range of interaction energies (and corresponding r_0 values) that can be realized in an electron-hole plasma in CQWs. We find that the exponential tails of the lines at low and medium powers merge with the high-power spectra. Indeed, one can reproduce the essential features of the observed behavior by calculating the interaction of a recombining dipole with the surrounding gas, taking into account the change of Γ with energy (Fig. 1D). Hence, this low-energy tail can be viewed as the spectral signature of interacting electron-hole plasma in CQWs. More details are given in section 1.1 of (9).

Fig. 1. The evolution of the photoluminescence spectrum with power. (A)

The energy band diagram of the asymmetric CQW under a perpendicular applied electric field. X_{WW} is the exciton in the wide well, X_{NW} is the exciton in the narrow well, and IX is the interwell indirect recombination. The blue circles represent electrons and the red circles represent holes. **(B)** Photoluminescence peak positions for X_{WW} , IX , and Z lines as a function of laser power at $T = 1.5$ K. **(C)** The evolution of the spectrum with power between 3 and 10 μW (the small peak at 1522 meV is the trion). The inset shows the spectra on a logarithmic scale. a.u., arbitrary units. **(D)** Calculated spectra demonstrating the formation of the low-energy exponential tail.



Department of Condensed Matter Physics, Weizmann Institute of Science, Rehovot, Israel.

*Corresponding author. E-mail: michael.stern@cea.fr

†Present address: Quantronics Group, SPEC, IRAMIS, DSM, CEA Saclay, Gif-sur-Yvette, France.

At higher power, above $P = 210 \mu\text{W}$, a new line, denoted by Z , appears below the IX line and shifts only slightly with power. One can show that this line is also due to the indirect recombination of electrons in the narrow well with holes in the wide

well. It shifts linearly with the gate voltage, approximately parallel to the IX line, remaining 3 to 4 meV above E_{IX}^0 , the IX energy at low power, $<1 \mu\text{W}$.

We found that the Z line came from a well-defined spatial region in the illuminated area. We

imaged the photoluminescence from the mesa and observed that below threshold, the image is smooth and follows the Gaussian profile of the excitation spot. Above threshold, a striking behavior appears: The luminescence separates into two regions, with a narrow (resolution-limited) dark border between them (Fig. 2A). The location of this boundary is unrelated to any structural parameter of the sample and can be affected by slight shifts of the exciting laser spot (movies S1 and S2).

The spectral properties of the two regions were analyzed by local photoluminescence measurements, with a $5\text{-}\mu\text{m}$ spatial resolution (Fig. 2B). As we moved across the boundary, the spectrum changed from being dominated by IX in the region denoted as I in Fig. 2A to being dominated by Z in region II.

In the following, we show that the Z line is a manifestation of an exciton liquid.

1) Criticality. The transition of the system into the phase-separated state is thermodynamic and occurs below a critical temperature and above a threshold power. Figure 2C shows the relative area of region II as a function of temperature (T). This phase appears abruptly at temperature $T = 4.7 \text{ K}$. Figure 2D demonstrates the power dependence: The appearance of region II above a threshold power is clearly seen. In this case, the area of region II extends gradually until it covers the entire sample.

2) High density. Time-resolved measurements reveal that the lifetime of the carriers in region II is ~ 60 to 80 ns (Fig. 2E), which is approximately twice as large as in region I. The phase separation therefore manifests two density regions: region I with lower density and region II with higher density. Using the absorption value of the CQW (12), we can estimate the density in region II to be $\sim 4 \times 10^{10}$ to $6 \times 10^{10} \text{ cm}^{-2}$ (9). At this high density, we get $n\rho^2 \approx 10$, fulfilling the

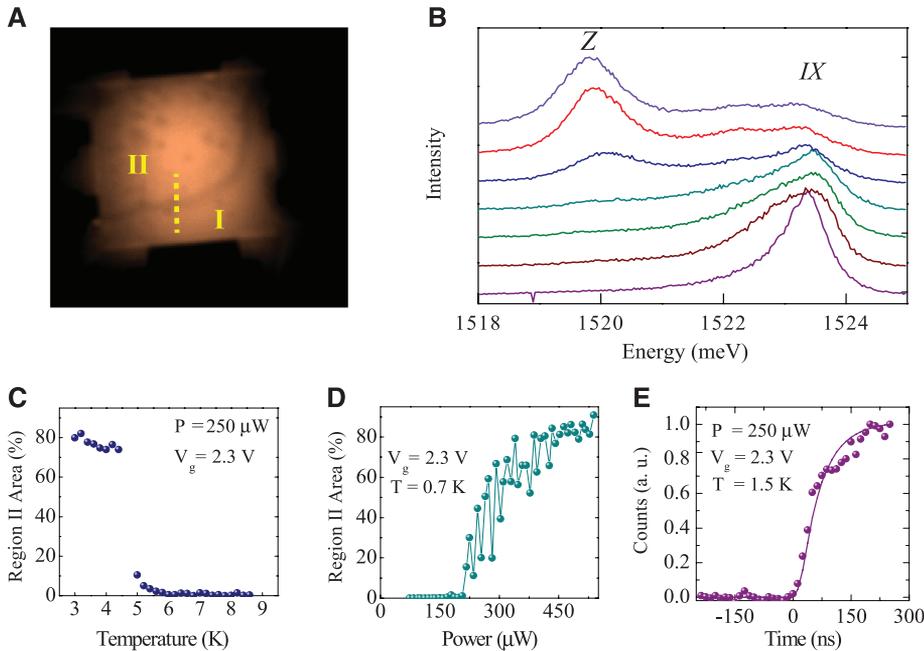


Fig. 2. The phase separation. (A) Image of the photoluminescence from the mesa showing the phase separation between region I and region II. The dark spots correspond to defects in the sample, which give reduced luminescence intensity in region II. This measurement was performed at $T = 1.5 \text{ K}$, $V_g = 2.3 \text{ V}$, and $P = 250 \mu\text{W}$. (B) Spectra measured along the dotted yellow line in (A) in the vicinity of the phase boundary. Each spectrum is shifted vertically by a fixed value, corresponding to its location along the line. (C and D) Evolution of the area of region II as a function of temperature (C) and laser power (D). The noise in the measurements reflects the fact that at some power levels, especially near threshold, the phase boundary exhibits large temporal fluctuations. (E) The rise of the Z line as a function of time (t) measured by correlated photon counting using an acousto-optic modulator. At $t = 0$, the laser is switched on and persists for $\sim 4 \mu\text{s}$. The solid line shows a fit to the experimental data, which gives a lifetime of $70 \pm 10 \text{ ns}$.

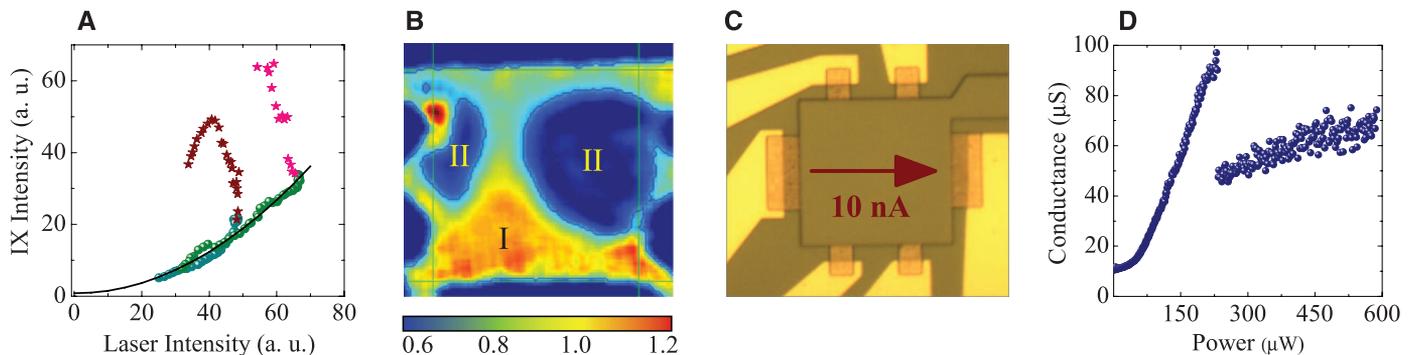


Fig. 3. The low diffusivity of region II. (A) A measurement of the local IX intensity as a function of the local laser intensity along two lines, similar to the dotted line in Fig. 2A. The brown and pink stars were measured in region I; the blue and green dots were measured in region II. The parabolic solid line demonstrates the correlation between the IX and laser intensities in region II. (B) The ratio between the photoluminescence images taken before and after phase separation. The thin green lines indicate the position of the mesa and are guides for the eyes. The blue color shows a reduction of intensity (<0.75)

after phase separation, whereas the red color represents an increase (>1.1). The diffusion away from the phase separation line toward the edges of the sample is clearly seen in region I. (C) Optical microscope image of the sample showing the four-probe electrical contacts to the $100\text{-}\mu\text{m}$ -square mesa. (D) Four-probe conductance as a function of laser power at $T = 1.5 \text{ K}$, $V_g = 2.3 \text{ V}$. The abruptness of the drop is due to the fact that the time interval between the measured points was 1 s , which is shorter than the nucleation time at $200 \mu\text{W}$ (9).

condition for a classical liquid, $n\rho^2 \gg 1$. The Z line energy is shifted relative to E_{IX}^0 by $\delta E = 3$ to 4 meV only. Using Eq. 1, we find that this corresponds to $r_0 \approx 25$ to 35 nm, which is approximately half the average interparticle distance at these densities.

3) Low diffusivity. Figure 3A shows the local intensity of the IX line as a function of the local excitation power. The IX intensity is well correlated with the excitation power in region II and is uncorrelated with it in region I. This implies that in region I, the carriers can recombine away from where they were excited, whereas in region II, they are localized. The high diffusivity of the carriers in region I is also seen when we compare the luminescence intensity profile in this region before and after the phase separation (Fig. 3B). The carriers in region I are pushed away from the region boundary and aggregate at the sides of the mesa.

The localization of carriers in region II is corroborated by in-plane electrical transport measurements conducted in a four-probe geometry (Fig. 3C). Figure 3D shows the conductance of the sample, σ , as a function of power. The onset of phase separation is clearly manifested by a drop of the conductance value and a reduction of the slope of the curve. Above the threshold power, the sample is divided into two regions, each with a different conductance, with $\sigma_I \gg \sigma_{II}$ (see materials and methods in the supplementary materials). The existence of a high-density region with low conductance is consistent with an exciton liquid.

4) Ordering. An important insight comes from studying the line shape of the Z line (Fig. 4A). The characteristic low-energy tail of the electron-hole plasma is absent, and the line is an almost perfect Gaussian (Fig. 4B). Noting that the low-energy tail is a manifestation of multiple values of r_0 , which characterize the gas phase, we are led to conclude that the Gaussian shape represents a narrow distribution around a dominant configuration. We used a local photoluminescence

measurement to extract the Z line at various locations and study its linewidth dependence on density. We found that the width of this Gaussian line decreases with increasing density (Fig. 4C). This line's narrowing cannot be explained by an uncorrelated gas, which should cause the linewidth to grow with density (13). To explain this behavior, we conducted a simple Monte Carlo calculation, assuming a fixed depletion region of radius r_0 around each exciton. With this assumption, the calculation reduces to randomly inserting N dipoles on a grid with r_0 periodicity and calculating their interaction energy. We found that the linewidth is indeed Gaussian and decreases with increasing density when $n\rho r_0^2 \geq 1/2$. Above this density, the number of possible configurations starts to decrease, giving rise to a narrower range of interaction energies. In fact, by taking $r_0 \approx 25$ nm, the value obtained from the estimated density, we get a quantitative agreement with the observed behavior (Fig. 4C).

The accumulation of these findings indicates that a strongly correlated ordered phase appears below a critical temperature and above a threshold density. The sample separates into two phases: an interacting electron-hole plasma in region I and a liquid phase in region II. Region II should not be perceived as a continuous liquid phase. The fact that we observe both Z and IX lines in region II (Fig. 2B) indicates that this region consists of a mixture of droplets and gas. Indeed, the existence of such a mixture was predicted by calculating the free energy of an exciton gas and solid (14).

Several other experimental findings are consistent with this interpretation. The first is the existence of a dark boundary between the two phases (Fig. 2A). This unusual behavior is explained by the fact that the carriers in the gas region experience a strong repulsive force from the localized high-density dipoles at the boundary. This repulsion creates a depletion region at their interface and induces a current of electron-

hole pairs toward the sides of the mesa (Fig. 3B). Another item of supporting evidence is related to the phase separation dynamics. We found clear evidence for a nucleation process: There is a relatively long (seconds) delay time between the onset of the laser light and the appearance of the phase separation (fig. S3). After this delay time, region II spreads from a corner of the mesa to its full extent (movie S3).

In closing, we wish to point out an unexplained observation. By comparing the integrated photoluminescence intensity before and after the phase separation, we find that the recombination in region II is not entirely radiative, and approximately 15% of the intensity is lost. Furthermore, we find that this darkening disappears at a weak perpendicular magnetic field, $B \sim 0.3$ T. Above this field, the phase separation is still observed. The fact that the phase separation can be observed without darkening rules out the formation of a dark condensate as the source of the observed behavior (8, 15). Further studies are needed to examine this behavior.

References and Notes

1. L. V. Keldysh, A. N. Kozlov, *Sov. Phys. JETP* **27**, 521 (1968).
2. X. Zhu, P. B. Littlewood, M. S. Hybertsen, T. M. Rice, *Phys. Rev. Lett.* **74**, 1633–1636 (1995).
3. Y. E. Lozovik, O. L. Berman, *JETP Lett.* **64**, 573–579 (1996).
4. B. Laikhtman, R. Rapaport, *Phys. Rev. B* **80**, 195313 (2009).
5. M. Stern, V. Garmider, V. Umansky, I. Bar-Joseph, *Phys. Rev. Lett.* **100**, 256402 (2008).
6. A. V. Gorbunov, V. B. Timofeev, *JETP Lett.* **96**, 138–147 (2012).
7. A. A. High *et al.*, *Nature* **483**, 584–588 (2012).
8. Y. Shilo *et al.*, *Nature Comm.* **4**, 2335 (2013).
9. See the supplementary materials on Science Online.
10. A. Alexandrou *et al.*, *Phys. Rev. B* **42**, 9225–9228 (1990).
11. G. J. Schinner *et al.*, *Phys. Rev. Lett.* **110**, 127403 (2013).
12. W. T. Masselink *et al.*, *Phys. Rev. B* **32**, 8027–8034 (1985).
13. Z. Vörös, D. W. Snoke, L. Pfeiffer, K. West, *Phys. Rev. Lett.* **103**, 016403 (2009).
14. R. Suris, R. Sergeev, paper presented at the Low Dimensional Structures Conference, Genoa, 2007.
15. M. Combescot, O. Betbeder-Matibet, R. Combescot, *Phys. Rev. Lett.* **99**, 176403 (2007).

Acknowledgments: We thank R. Suris, B. Laikhtman, M. and R. Combescot, and V. Steinberg for fruitful discussions and M. Brook for the magnetic field measurements. This work was supported by the Israeli Science Foundation.

Supplementary Materials

www.sciencemag.org/content/343/6166/55/suppl/DC1
Materials and Methods
Supplementary Text
Figs. S1 to S5
References (16–18)
Movies S1 to S3

18 July 2013; accepted 25 November 2013
10.1126/science.1243409

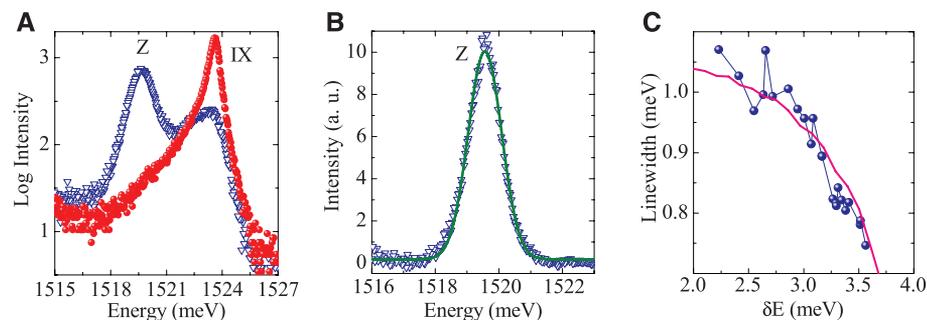


Fig. 4. Shape and width of the Z line. (A) A comparison between the photoluminescence spectrum in region I (red) and region II (blue). (B) The spectrum of the Z line obtained after removal of the exponential energy dependence, $\Gamma(E)$. The solid green line shows an excellent fit to a Gaussian. (C) The width of the Z line as a function of the energy shift δE . The reduction of the line width with density is clearly seen. The solid red line shows the width obtained by Monte Carlo simulations, taking into account 0.4 meV of inhomogeneous broadening (half width at half maximum). A good agreement between measurement and calculation is obtained for $r_0 = 25$ nm.